## Physics cup 2024 problem 1

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There are two methods of solving this problem, the first one is done by solving Laplace's equation, and the second one is done by making an analogy to the magnetic field around a superconductor. While I have used the first method for my initial submission, I think it is important to include the second one as it is less complex mathematically.

## Method 1:

First, to describe the motion of water around the disk, I did the following:

## Figuring out the motion of the water

Water, as we know, is incompressible. So for each volume element of it, the water that gets in a $d x^{*} d y^{*} d z$ volume is the same as the water that gets out

$\left(v_{y_{(y)}}-v_{y_{y+d y}}\right) d x d z+\left(v_{x_{(x)}}-v_{x_{x+d x}}\right) d y d z+\left(v_{z_{(z)}}-v_{z_{z+d z}}\right) d x d y=0$
$\frac{d v_{y}}{d y}+\frac{d v_{x}}{d x}+\frac{d v_{z}}{d z}=0$
Let's define a velocity potential, $\phi$, so
$\frac{d \phi}{d y}=v_{y}, \frac{d \phi}{d x}=v_{x}, \frac{d \phi}{d z}=v_{z}$
Therefore:
$\frac{d^{2} \phi}{d y^{2}}+\frac{d^{2} \phi}{d x^{2}}+\frac{d^{2} \phi}{d z^{2}}=0$
In order to find the potential flow around the disk, we will first look at the general case of flow around an ellipsoid, and make the length of its main axis approach 0.

the flow around an ellipse
We would first start by establishing a 3d elliptical coordinate system, then establish boundary conditions and solve the laplace equation.
Since a spheroid is the rotation shape of an ellipse, let's make this a 2d elliptic coordinate system with the third dimension being rotation around the y axis.

a 2D elliptical coordinate system, blue lines are of constant $v$ (the $\eta$ axis) and red lines are of constant $\eta$ (the $v$ axis). Made with desmos

elliptical coordinate system, from Wikipedia
Elliptical coordinates are defined such as when the axis $\eta$ is at a constant value, $x$ and $y$ create a ellipses, and when the axis $v$ is at a constant value, $x$ and $y$ create hyperbolae with the same focal points
$x_{0}=k * \sinh (\eta) \sin (v)$
$y_{0}=k * \cosh (\eta) \cos (v)$
where k is a constant of scale
so
$\frac{x^{2}}{\sinh ^{2}(\eta)}+\frac{y^{2}}{\cosh ^{2}(\eta)}=\sin ^{2}(v)+\cos ^{2}(v)=1$
and
$\frac{x^{2}}{\sin ^{2}(v)}-\frac{y^{2}}{\cos ^{2}(v)}=\sinh ^{2}(\eta)-\cosh ^{2}(\eta)=1$
Introducing an angle $\omega$ from the y to z axis, and to define the ellipsoid we would convert $x_{0}$ from the previous system to the distance from the x axis $r=x_{0}=x_{0} \cos (\varphi) \hat{x}_{1}+x_{0} \sin (\omega) \hat{z}_{1}$ $y_{1}=k * \cosh (\eta) \sin (v) \cos (\omega)$
$x_{1}=k * \sinh (\eta) \cos (v)$
$z_{1}=k^{*} \cosh (\eta) \sin (v) \sin (\omega)$
we simplify according to $\sin ^{2}(\nu)+\cos ^{2}(\nu)=1$, and $\sinh ^{2}(\mu)-\cosh ^{2}(\mu)=-1$
$y_{1}=k \sqrt{1+\zeta^{2}} \sqrt{1-\mu^{2}} \cos (\omega)$
$x_{1}=k \zeta \mu$
$z_{1}=k \sqrt{1+\zeta^{2}} \sqrt{1-\mu^{2}} \sin (\omega)$

in the figure, an ellipsoid defined by $\frac{x^{2}}{\sinh ^{2}(\eta)}+\frac{y^{2}+z^{2}}{\cosh ^{2}(\eta)}=1$, made with desmos 3d
We now define a length factor for each of our new coordinates, the length factor is defined as the distance from two points at a curtain place in our coordinate system

$\frac{d s_{\zeta}}{d \zeta}=\sqrt{\left(\frac{d x}{d \zeta}\right)^{2}+\left(\frac{d y}{d \zeta}\right)^{2}+\left(\frac{d z}{d \zeta}\right)^{2}}=k \sqrt{\left(\frac{\zeta \sqrt{1-\mu^{2}}}{\sqrt{1+\zeta^{2}}} \cos (\omega)\right)^{2}+(\mu)^{2}+\left(\frac{\zeta \sqrt{1-\mu^{2}}}{\sqrt{1+\zeta^{2}}} \sin (\omega)\right)^{2}}$
$=k \sqrt{\left(\frac{\zeta^{2}\left(1-\mu^{2}\right)+\mu^{2}\left(1+\zeta^{2}\right)}{1+\zeta^{2}}\right.}=k \sqrt{\frac{\zeta^{2}+\mu^{2}}{\zeta^{2}-1}}$
$\frac{d s_{\mu}}{d \mu}=\sqrt{\left(\frac{d x}{d \mu}\right)^{2}+\left(\frac{d y}{d \mu}\right)^{2}+\left(\frac{d z}{d \mu}\right)^{2}}=k \sqrt{\left(\frac{-\mu \sqrt{1+\zeta^{2}}}{\sqrt{1-\mu^{2}}} \cos (\omega)\right)^{2}+(\zeta)^{2}+\left(\frac{-\mu \sqrt{1+\zeta^{2}}}{\sqrt{1-\mu^{2}}} \sin (\omega)\right)^{2}}=$
$k \sqrt{\frac{\mu^{2}\left(1+\zeta^{2}\right)+(\zeta)^{2}\left(1-\mu^{2}\right)}{1-\mu^{2}}}=k \sqrt{\frac{\zeta^{2}+\mu^{2}}{1-\mu^{2}}}$
$\frac{d s_{\omega}}{d \omega}=\sqrt{\left(\frac{d x}{d \omega}\right)^{2}+\left(\frac{d y}{d \omega}\right)^{2}+\left(\frac{d z}{d \omega}\right)^{2}}=k \sqrt{\left(\sqrt{1+\zeta^{2}} \sqrt{1-\mu^{2}} \sin (\omega)\right)^{2}+\left(\sqrt{1+\zeta^{2}} \sqrt{1-\mu^{2}} \cos (\omega)\right)^{2}}=$
$k \sqrt{\left(1+\zeta^{2}\right)\left(1-\mu^{2}\right)}=\sqrt{\zeta^{2}+1} \sqrt{1-\mu^{2}}$
using these elements, we will define the new velocity potential $\phi$ as
$\frac{d \phi}{d s_{\mu}}=v_{\mu} \quad, \frac{d \phi}{d s_{\zeta}}=v_{\zeta} \quad, \frac{d \phi}{d s_{\omega}}=v_{\omega}$

Now we can do the same thing we did in part 1, but each area element will be the product of two length elements,
$\frac{d}{d \mu}\left(\frac{d s_{\omega}}{d \omega} \frac{d s_{\zeta}}{d \zeta} \frac{d \phi}{d s_{\mu}}\right)+\frac{d}{d \zeta}\left(\frac{d s_{\omega}}{d \omega} \frac{d s_{\mu}}{d \mu} \frac{d \phi}{d s_{\zeta}}\right)+\frac{d}{d \omega}\left(\frac{d s_{\zeta}}{d \zeta} \frac{d s_{\mu}}{d \mu} \frac{d \phi}{d s_{\omega}}\right)=0$
Because of symmetry, we know $v_{\omega}=0$
$\frac{d}{d \mu}\left(\frac{d s_{\omega}}{d \omega} \frac{d s_{\zeta}}{d \zeta} \frac{d \mu}{d s_{\mu}} \frac{d \phi}{d \mu}\right)+\frac{d}{d \zeta}\left(\frac{d s_{\omega}}{d \omega} \frac{d s_{\mu}}{d \mu} \frac{d \zeta}{d s_{\zeta}} \frac{d \phi}{d \zeta}\right)=0$
$\frac{d}{d \mu}\left(\left(1-\mu^{2}\right) \frac{d \phi}{d \mu}\right)+\frac{d}{d \zeta}\left(\left(\zeta^{2}+1\right) \frac{d \phi}{d \zeta}\right)=0$
we will solve this with separation of variables
$\phi=u k A_{(\zeta)} B_{(\mu)}$
$\frac{\frac{d}{d \mu}\left(\left(1-\mu^{2}\right) \frac{d B}{d \mu}\right)}{B}=c$
$c=\alpha(\alpha+1)$
$\frac{d}{d \mu}\left(\left(\mu^{2}-1\right) \frac{d B}{d \mu}\right)=\alpha(\alpha+1) B$
$B=\sum_{n=-\infty}^{\infty} a_{n} \mu^{n}$
$\frac{d}{d \mu}\left(\sum_{n=-\infty}^{\infty} n a_{n} \mu^{n+1}-\sum_{n=-\infty}^{\infty} n a_{n} \mu^{n-1}\right)=\alpha(\alpha+1) \sum_{n=-\infty}^{\infty} a_{n} \mu^{n}$
$\sum_{n=-\infty}^{\infty}(n+1) n a_{n} \mu^{n}-\sum_{n=-\infty}^{\infty} n(n-1) a_{n} \mu^{n-2}=\alpha(\alpha+1) \sum_{n=-\infty}^{\infty} a_{n} \mu^{n}$
$a_{0}=a_{0}$
$a_{0} \alpha(\alpha-1)=-2 a_{2}$
$a_{1} \alpha(\alpha-1)=2 a_{1}-6 a_{3}$
$a_{n-2}(-\alpha(\alpha-1)+(n-2)(n-1))=a_{n}(n)(n-1)$
$a_{n-2}=\frac{a_{n}(n)(n-1)}{(n-2)(n-1)-\alpha(\alpha+1)}$
$a_{n+2}=\frac{a_{n}(n(n+1)-\alpha(\alpha+1))}{(n+2)(n+1)}$
The solution to this type of equation is a well known one, this is Legendre's differential equation, and it can be solved with the Legendre functions of the first and second kind $B=\left(a \cdot P_{\alpha}(\mu)+b \cdot Q_{\alpha}(\mu)\right)$
These unique functions are used everywhere from quantum mechanics to Mie scattering.

We know that when $\zeta \rightarrow \infty$, the velocity towards the $\mu$ direction in an almost straight angle ( $\mu \rightarrow 1$ ) the velocity of the water is $u$ to the $x$ direction, therefore it is $u$ times $\frac{\sqrt{y^{2}+z^{2}}}{\sqrt{x^{2}+y^{2}+z^{2}}} \approx_{\zeta \rightarrow \infty} \sqrt{1-\mu^{2}}$, to the $\mu$ direction. we also know because of that when $\zeta \rightarrow \infty$, the velocity doesn't jump up to infinity, therefore the velocity potential function is only limited to the first order, so $\alpha=1$
$a_{n+2}=\frac{a_{n}(n(n+1)-2)}{(n+2)(n+1)}=\frac{a_{n}(n+2)(n-1)}{(n+2)(n+1)}=\frac{a_{n}(n-1)}{(n+1)}$
$P_{1}(x)=a_{0} x$
$a_{-n-2}=\frac{a_{n}(-n-1)}{(-n-3)}=\frac{a_{n+2}(-n+1)}{(-n-3)}=\ldots=\frac{a_{0}(0+1)}{(-n-3)}$ or $\frac{a_{1}(-1+1)}{(-n-3)}$
for $\mathrm{n}<0$
$a_{-2 n}=\frac{a_{0}}{(-2 n-1)}$
$a_{-2 n-1}=0$
for $\mathrm{n}=0$ :
$a_{n=0}=\frac{a_{0}}{(0-1)}=-a_{n=0}$
$a_{n=0}=0$
$a_{-n-2}=\frac{a_{-n}(-n-1)}{(-n-3)}=\frac{a_{-n+2}(-n+1)}{(-n-3)}=\ldots=\frac{-a_{-2}}{(-n-3)}$ or $\frac{a_{-1}(-1+1)}{(-n-3)}$
$Q_{1}(x)=-\sum_{n=1}^{\infty} \frac{a_{-2}}{x^{2 n}(2 n+1)}$
Let's write our boundary condition for $\mu$
$\frac{d \phi}{d s_{\mu}}=\frac{\frac{d \phi}{d s_{\mu}}}{d_{\mu=1, \zeta \rightarrow \infty}} \underset{k \sqrt{\frac{z^{2}+\mu^{2}}{1-\mu^{2}}}}{ }=u \sin (v)=u \sqrt{1-\mu^{2}}$
$\left(a+b \sum_{n=1}^{\infty} \frac{-2 n^{*} \mu^{-2 n-1}}{(2 n+1)}\right) \sqrt{1-\mu^{2}}=$ const $* \sqrt{1-\mu^{2}}$
$(a-\infty b)=$ const
$b=0$
$B=a \mu$
For the function $A$, the solution is similar, we just need to add the imaginary unit to $\zeta$ $A=\left(c \cdot P_{1}(i \zeta)+d \cdot Q_{1}(i \zeta)\right)=$
$\left(c \cdot i \zeta+d \cdot\left(-\sum_{n=1}^{\infty} \frac{a_{-2}}{(i \zeta)^{2 n}(2 n+1)}\right)\right)=\left(c \cdot i \zeta+d \cdot\left(a_{-2}-\sum_{n=0}^{\infty} \frac{a_{-2}}{(i \zeta)^{2 n}(2 n+1)}\right)\right)=$
$\left(c \cdot i \zeta-d \cdot\left(\zeta i\left(\sum_{n=0}^{\infty} \frac{a_{-2}}{(i \zeta)^{2 n+1}(2 n+1)}\right)-a_{-2}\right)\right)$
$\left(c \cdot i \zeta-d \cdot\left(i \zeta\left(\left(\frac{a_{-2}}{(i \zeta)(1)}\right)+\left(\frac{a_{-2}}{\left(-i \zeta^{3}\right)(3)}\right)+\left(\frac{a_{-2}}{\left(i \zeta^{5}\right)(5)}\right)+\left(\frac{a_{-2}}{\left(-i \zeta^{7}\right)(7)}\right) \ldots.\right)-a_{-2}\right)\right)=$
$\left(c \cdot i \zeta-d \cdot a_{-2}\left(\zeta\left(\left(\frac{1}{(\zeta)(1)}\right)-\left(\frac{1}{\left(\zeta^{3}\right)(3)}\right)+\left(\frac{1}{\left(\zeta^{5}\right)(5)}\right)-\left(\frac{1}{\left(\zeta^{7}\right)(7)}\right) \ldots\right)-1\right)\right)=$
$\left(c \cdot i \zeta-d \cdot a_{-2}\left(\zeta \int\left(\left(\frac{1}{\left(\zeta^{2}\right)}\right)-\left(\frac{1}{\left(\zeta^{4}\right)}\right)+\left(\frac{1}{\left(\zeta^{6}\right)}\right)-\left(\frac{1}{\left(\zeta^{8}\right)}\right) \ldots\right) d\left(\frac{1}{\zeta}\right)-1\right)\right)=$
$\left(c \cdot i \zeta-d \cdot a_{-2}\left(\zeta \int\left(\frac{\frac{1}{\zeta^{2}}}{1+\frac{1}{\zeta^{2}}}\right) d\left(\frac{1}{\zeta}\right)-1\right)\right)=$
$\left(c \cdot i \zeta-d \cdot a_{-2} \cdot\left(\zeta \tan ^{-1}\left(\frac{1}{\zeta}\right)-1\right)\right)=\left(c \cdot i \zeta+d \cdot\left(1-\zeta \cot ^{-1}(\zeta)\right)\right)$
To find c and d, we know that when $\zeta \rightarrow \infty$, the velocity towards the $\zeta$ direction is $u$ times.
$\frac{x}{\sqrt{x^{2}+y^{2}+z^{2}}} \approx_{\zeta \rightarrow \infty} \mu$,
${\frac{d \phi}{d s_{\zeta}}}_{\mu \rightarrow 1, \zeta \rightarrow \infty}=u \mu$
$u k \frac{d \phi}{d \zeta} \frac{1}{k} \sqrt{\frac{\zeta^{2}-1}{\zeta^{2}+\mu^{2}}}{ }_{\mu \rightarrow 1, \zeta \rightarrow \infty}=u \mu$
$u a \mu \cdot \sqrt{\frac{\zeta^{2}-1}{\zeta^{2}+1}} \cdot\left(c \cdot i \zeta+d \cdot\left(1-\zeta \cot ^{-1}(\zeta)\right)=u \mu\right.$
$a \cdot(c \cdot i+d \cdot(1-1))=1$
$c=\frac{1}{a i}$
$A=d \cdot\left(1-\zeta \cdot \cot ^{-1}(\zeta)\right)+\frac{1}{a i} \cdot i \zeta$
$\phi=u k \mu a\left(d \cdot\left(1-\zeta \cdot \cot ^{-1}(\zeta)\right)+\frac{1}{a} \zeta\right)=u k \mu\left(g_{1} \cdot\left(\zeta \cdot \cot ^{-1}(\zeta)-1\right)+\zeta\right), g_{1}=-d a$
*** note: the choice of variables is not meant to endorse or make fun of the United Kingdom, any mention of the letters $u k$ is coincidental only.

Now let's find the parameter $g_{1}$ :
We know that the points where $\zeta=\zeta_{0}$, form an ellipsoid in our 3D space. To describe the flow around it we can impose a boundary condition that no water at the surface of the ellipsoid moves perpendicular to it.

$$
\begin{aligned}
& \frac{d \phi}{d s_{\zeta}}=0 \\
& \frac{d \phi}{d \zeta} \frac{1}{k} \sqrt{\frac{\zeta_{0}^{2}-1}{\zeta^{2}+\mu^{2}}}=0 \\
& \frac{d \phi}{d \zeta} \zeta_{\zeta=\zeta_{0}} \frac{1}{k} \sqrt{\frac{\zeta_{0}^{2}-1}{\zeta_{0}^{2}+\mu^{2}}}=0 \\
& u k \mu\left(g_{1} \cdot\left(\cot ^{-1}(\zeta)-\frac{\zeta}{\zeta^{2}+1}\right)+1\right)_{\zeta=\zeta_{0}}^{\frac{1}{k}} \sqrt{\frac{\zeta_{0}^{2}-1}{\zeta_{0}^{2}+\mu^{2}}}=0 \\
& \left(g_{1} \cdot\left(\cot ^{-1}\left(\zeta_{0}\right)-\frac{\zeta \zeta_{0}}{\zeta_{0}^{2}+1}\right)+1\right) \sqrt{\frac{\zeta_{0}^{2}-1}{\zeta_{0}^{2}+\mu^{2}}}=0 \\
& g_{1} \cdot\left(\cot ^{-1}\left(\zeta_{0}\right)-\frac{\zeta_{0}}{\zeta_{0}^{2}+1}\right)=-1
\end{aligned}
$$

$g_{1}=\frac{1}{\left(\frac{\zeta_{0}}{\zeta_{0}^{2}+1}-\cot ^{-1}\left(\zeta_{0}\right)\right)}$

If we choose now to make $\zeta_{0} \rightarrow 0$, our new ellipsoid will now become:
$\frac{x^{2}}{\sinh ^{2}(\eta)}+\frac{y^{2}+z^{2}}{\cosh ^{2}(\eta)}=k^{2}$
$\frac{x^{2}}{\zeta_{0}^{2}}+\frac{y^{2}+z^{2}}{\zeta_{0}^{2}+1}=k^{2}$
$x_{\max , \min }= \pm k \zeta_{0} \rightarrow 0$
$y^{2}+z^{2}<k^{2}$
we got a disk with radius k .
Now let's push $g_{1}$ to the limit where $\zeta_{0} \rightarrow 0$
$g_{1}=\frac{1}{\left(\frac{0}{0^{2}+1}-\cot ^{-1}(0)\right)}=\frac{-2}{\pi}$
$\phi=u k \mu\left(\frac{2}{\pi}\left(1-\zeta \cdot \cot ^{-1}(\zeta)\right)+\zeta\right)$

In the lab's frame of reference, the water far away is stationary and only the water near the disk moves. To move to that reference frame we need to subtract a uniform flow in the $x$ direction from our potential function:
$\phi_{\text {uniform }}=x u=u k \zeta \mu$
$\phi=u k \mu\left(\frac{2}{\pi}\left(1-\zeta \cdot \cot ^{-1}(\zeta)\right)\right)$
Now, to find the force enacted on the plate to the $x$ direction, we will integrate the pressure over the surface of the disk, the pressure is defined as
$P=\rho\left(\frac{d \phi}{d t}+\frac{1}{2} \nabla \phi^{2}\right)$
at the moment of impact with the water, the change in velocity (and thus in potential) is almost instantaneous, therefore we can neglect any other term in the
$F=-\rho \int_{0}^{2 \pi} \int \frac{d \phi}{d t} d \zeta_{\mu} d s_{\omega}=$
$-\rho \dot{u} k^{3} \iint \mu\left(\frac{2}{\pi}\left(1-\zeta_{0} \cdot \cot ^{-1}\left(\zeta_{0}\right)\right)\right) \sqrt{\frac{\zeta_{0}^{2}+\mu^{2}}{1-\mu^{2}}} \sqrt{\zeta_{0}^{2}+1} \sqrt{1-\mu^{2}} d \mu d \omega=$
$-\rho u k^{3} \int_{0}^{2 \pi} \int \mu\left(\frac{2}{\pi}\left(1-\zeta_{0} \cdot \cot ^{-1}\left(\zeta_{0}\right)\right)\right) \sqrt{\zeta_{0}^{2}+\mu^{2}} \sqrt{\zeta_{0}^{2}+1} d \mu d \omega=$
$-\rho \dot{u} k^{3} \int_{0}^{2 \pi} \int \mu\left(\frac{2}{\pi}(1)\right) \sqrt{0^{2}+\mu^{2}} \sqrt{0^{2}+1} d \mu d \omega=$
$-\rho \dot{u k}{ }^{3} \int_{0}^{2 \pi} \int \frac{2}{\pi} \mu^{2} d \mu d \omega=$
$-2 \pi \cdot \frac{2}{\pi} \cdot \rho \dot{u} k^{3} \int \mu^{2} d \mu=$

Again, $\mu$ represents the cosine of the angle between the $x$ axis and the $y z$ plane when $\zeta \rightarrow \infty$ . Since we only care about one side of the disk, the limits of integration for the angle are 0 and $\frac{\pi}{2}$, or 0 and 1 for $\mu$.
$-4 \cdot \operatorname{\rho uk}{ }^{3}\left[{ }_{0}^{1} \frac{\mu^{3}}{3} d \mu=\right.$
$-\rho \frac{4}{3} \dot{u} k^{3}$
this can be seen as a mass of $M_{\text {added }}=\frac{4}{3} \rho k^{3}$ added to the cylinder in an inelastic collision, we can write the conservation of momentum equation

$$
\begin{aligned}
& \left(M_{c y l i n d e r}+M_{\text {added }}\right) \cdot u_{\text {after }}=M_{\text {cylinder }} \cdot u_{\text {before }} \\
& M_{\text {cylinder }}=\pi \rho_{\text {cylinder }} \cdot(k)^{2} \cdot H=\frac{\frac{4}{3} \rho k^{3} u_{\text {after }}}{u_{\text {before }}-u_{\text {after }}} \\
& \rho_{\text {cylinder }}=\frac{4 \rho k u_{\text {after }}}{3 \pi\left(u_{\text {before }}-u_{\text {after }} \cdot(H)\right.}
\end{aligned}
$$

## Method 2:

the Laplace equation, or $\frac{d^{2} \phi}{d y^{2}}+\frac{d^{2} \phi}{d x^{2}}+\frac{d^{2} \phi}{d z^{2}}=0$ is also present in electromagnetism. In the case of electromagnetism, the velocity field is equivalent to the magnetic field around a polarized object.

We can find the added mass by solving an auxiliary problem of figuring the field around an infinitely thin circular plate. Our boundary conditions are that no water gets inside the body. In other words, the field orthogonal to the surface at the surface is zero and no movement of water happens inside the body.
This is also the behavior of the magnetic field around a superconducting object. In such an object, the surface currents arrange in such a way that the magnetic field inside is zero, therefore, if the applied field is uniform, the magnetic field created by the object inside of it will also be uniform.
We would treat the plate as an infinitely thin ellipsoid. The reason for that is that when the magnetization of it is uniform, the magnetic field caused by said magnetization is also uniform.
to prove this, we can use the electri-magnetic analogy,
we start with a homogeneously charged spherical shell, defined by the superposition of $x^{2}+y^{2}+z^{2}=R^{2}$ and $x^{2}+y^{2}+z^{2}=(R+d r)^{2}$
we pick a point P inside of it, and draw two narrow cones meeting at the point.

let's find the magnitude of the electric field at the point from both surfaces:
$E=\frac{k \rho d V_{1}}{h_{1}{ }^{2}}-\frac{k \rho d V_{2}}{h_{2}{ }^{2}}=\frac{k \sigma A_{1}}{h_{1}{ }^{2}}-\frac{k \sigma A_{2}}{h_{2}{ }^{2}}=\frac{k \sigma A_{1}}{h_{1}{ }^{2}}-\frac{k \sigma A_{2}}{h_{2}{ }^{2}}=\frac{k \sigma \frac{\left(\pi h_{1}{ }^{2} \alpha\right)}{\cos ^{2}(\theta)}}{h_{1}{ }^{2}}-\frac{k \sigma \frac{\left(\pi h_{2}{ }^{2} \alpha\right)}{\cos ^{2}(\theta)}}{h_{2}{ }^{2}}=0$
we can sum these cones over the entire surface of the sphere and get that at every point $E=0$

Now, we pick a point $\mathrm{P}^{\prime}$ inside a homogeneously charged ellipsoidal shell, the shell is defined by the superposition of $\frac{x^{2}}{b^{2}}+y^{2}+z^{2}=R^{2}$ and $\frac{x^{2}}{b^{2}}+y^{2}+z^{2}=(R+d R)^{2}$. we can solve this by applying an affine transformation from the spherical case of $x \rightarrow \frac{x}{b}$

$d V_{1}=d x_{1} d y_{1} d z_{1} \rightarrow d z_{1} d y_{1} \frac{d x_{1}}{b}$
$d V_{2}=d x_{2} d y_{2} d z_{2} \rightarrow d z_{2} d y_{2} \frac{d x_{2}}{b}$

We also notice that the translation of the two bases of the cone in relation to each other, as well as the translation of the point $P$ are parallel to the $x$ axis, moving the two centers to align, and we get two intersecting lines and two parallel lines creating two triangles, which we know from the Talos principle:
$\frac{h_{1}}{h_{2}}=\frac{h_{1,}}{h_{2_{f}}}$
therefore:
$E=\frac{k \rho d V_{1_{f}}}{h_{1,}{ }^{2}}-\frac{k \rho d V_{2_{f}}}{h_{2_{f}}{ }^{2}}=\frac{k \rho d V_{1}}{b h_{1, f}{ }^{2}}-\frac{k \rho d V_{2}}{b h_{2}{ }^{2}}=\frac{1}{h_{1,}{ }^{2}}\left(\frac{k \rho d V_{1}}{b}-\frac{k \rho d V_{2} h_{1}{ }^{2}}{b h_{2}{ }^{2}}\right)=\frac{h_{1}{ }^{2}}{b h_{1,}{ }^{2}}\left(\frac{k \rho d V_{1}}{h_{1}{ }^{2}}-\frac{k \rho d V_{2}}{h_{2}{ }^{2}{ }^{2}}\right)=0$ we now define a new size $\zeta$, each point has an ellipsoid of $\frac{x^{2}}{b^{2}}+y^{2}+z^{2}=\zeta^{2}$ touching it. By applying gauss's law for each layer of a uniformly charged ellipsoid, we find that the electric field inside of it is $E=c^{*} \zeta$ normal to the surface of the ellipsoid $\frac{x^{2}}{b^{2}}+y^{2}+z^{2}=\zeta^{2}$ where $c$ is a constant.
Therefore:
$E_{x}=c * \zeta \frac{d \zeta}{d x}=c * \sqrt{\frac{x^{2}}{b^{2}}+y^{2}+z^{2}} \frac{x}{b^{2} \sqrt{\frac{x^{2}}{b^{2}}+y^{2}+z^{2}}}=\frac{c x}{b^{2}}$
$E_{y}=c * \zeta \frac{d \zeta}{d y}=c * \sqrt{\frac{x^{2}}{b^{2}}+y^{2}+z^{2}} \frac{y}{\sqrt{\frac{x^{2}}{b^{2}}+y^{2}+z^{2}}}=c y$
$E_{z}=c * \zeta \frac{d \zeta}{d z}=c * \sqrt{\frac{x^{2}}{b^{2}}+y^{2}+z^{2}} \frac{z}{\sqrt{\frac{x^{2}}{b^{2}}+y^{2}+z^{2}}}=c z$
A uniformly polarized ellipsoid in the x direction is a superposition of a uniformly positively charged ellipsoid at $x=0$ and an oppositely charged ellipsoid at $x=d x$, so for each point of $\left(x_{0}, y_{0}, z_{0}\right)$ inside the ellipsoid:
$E_{x}=\frac{c\left(\left(x_{0}-0\right)-\left(x_{0}-d x\right)\right)}{b^{2}}=\frac{c^{*} d x}{b^{2}}, E_{y}=c\left(y_{0}-y_{0}\right)=0, E_{z}=c\left(z_{0}-z_{0}\right)=0$
Therefore a uniformly polarized ellipsoid creates a uniform electric field inside of it.
To apply the electric-magnetic analogy, we need to consider only fields outside the body. We can prove the magnetic field is homogeneous at an arbitrary point $A$ inside the ellipsoid by drilling a narrow hole parallel to the magnetization axis next to it, for the points inside the hole the analogy holds and the magnetic field is uniform.
For the point $A$ outside the hole, we know that the H field parallel to the surface is conserved, the H field is defined by the magnetic field plus $\mu_{0} M$ (which is constant), therefore Since the magnetic field inside the hole is uniform, then H is uniform and the magnetic field at point $A$ is also uniform. since this can be applied to any point in the ellipsoid, the magnetic field inside the entire ellipsoid is homogenous

We now need to find the polarization of the ellipsoid, we can do that by comparing the field produced by the polarized ellipsoid and the field exerted on it. It is known that the bound surface current is $K_{m}=M \times \hat{n}=M(\hat{x} \times \hat{n})$
imagine a cross section of the ellipsoid with the plane $z=0$, from symmetry the same will apply to any plane parallel to the $x$ axis and going through the center.

$\frac{x^{2}}{b^{2}}+y^{2}=R^{2}$
$x= \pm b \sqrt{R^{2}-y^{2}}$
the angle between the tangent of the ellipse and the y axis is $\varphi$ $\tan (\varphi)=\frac{d x}{d y}=\frac{\mp b y}{\sqrt{R^{2}-y^{2}}} \Rightarrow \sin (\varphi)=\frac{\tan (\varphi)}{\sqrt{1+\tan ^{2}(\varphi)}}=\frac{\mp b y}{\sqrt{b^{2} y^{2}+R^{2}-y^{2}}}$
due to rotational symmetry, the $\vec{n}$ vector at the points that lay on the perimeter of the ellipse has no $\hat{z}$ component. Therefore:
$|(\hat{x} \times \hat{n})|=\sin (\varphi)=\frac{\mp b y}{\sqrt{b^{2} y^{2}+R^{2}-y^{2}}}$
$K_{m}=\frac{\mp b y M}{\sqrt{b^{2} y^{2}+R^{2}-y^{2}}} \hat{z}$
Generalized to any plane parallel to the x axis and going through the center:
$x= \pm b \sqrt{R^{2}-r^{2}}, \vec{r}=\vec{y}+\vec{z}$
$\tan (\varphi)=\frac{\mp b r}{\sqrt{R^{2}-r^{2}}}$
$K_{m}=\frac{\mp b r M}{\sqrt{b^{2} r^{2}+R^{2}-r^{2}}}(\hat{x} \times \hat{r})$
To find the magnetic field, we will calculate the magnetic field at the center of the ellipsoid. Notice that the induced current on the surface is composed of current loops going around the center of the ellipsoid.
We now sum the magnetic fields generated by the loops, using the formula for the magnetic field generated by a ring at a point on its axis. Notice that the field applied from the top part is the same as the field in the bottom
$B=\frac{\mu_{0} I r^{2}}{2\left(r^{2}+x^{2}\right)^{3 / 2}}=2 \int_{0}^{R} \frac{\mu_{0} r^{2} K_{m} \cdot \frac{d r}{\cos (\varphi)}}{2\left(r^{2}+x^{2}\right)^{3 / 2}}=2 \int_{0}^{R} \frac{\mu_{0} r^{2} K_{m} \cdot \frac{d r}{\cos (\varphi)}}{2\left(r^{2}+x^{2}\right)^{3 / 2}}=2 \int_{0}^{R} \frac{\mu_{0} r^{2} M \tan (\varphi) \cdot d r}{2\left(r^{2}+x^{2}\right)^{3 / 2}}=2 \int_{0}^{R} \frac{\mu_{0} r^{2} M \frac{b r}{2\left(r^{2}+b^{2} R^{2}-b^{2}\right.} \cdot d r}{\left.2 r^{2}\right)^{3 / 2}}$
in the case of a flat disc $b \rightarrow 0$
$B=2 b \int_{0}^{R} \frac{\mu_{0} r^{3} M \cdot d r}{2 r^{3} \sqrt{R^{2}-r^{2}}}=b \mu_{0} M \int_{0}^{R} \frac{d\left(\frac{r}{R}\right)}{\sqrt{1-\left(\frac{r}{R}\right)^{2}}}=\frac{\pi}{2} b \mu_{0} M=\frac{\pi}{2} b \mu_{0} \frac{m}{b^{* \frac{4}{3} \pi R^{3}}}=\frac{3}{8} \mu_{0} \frac{m}{R^{3}}$
$m=\frac{8 B R^{3}}{3 \mu_{0}}$

To find the added mass, we need to find the energy of the moving water from the reference frame of the lab. Moving to the lab's frame of reference is the same as adding to the object's frame a field of constant velocity $-v$. Which is akin to subtracting the applied magnetic field and only summing the potential energy from the induced magnetic field.

We know that $U=\frac{1}{2} m B=\frac{4 B^{2} R^{3}}{3 \mu_{0}}$ which is equivalent to $E=\frac{1}{2} m v^{2}$ since $B$ is equivalent to $v$ , than $\frac{2 E}{B^{2}}=\frac{8 R^{3}}{3 \mu_{0}}$ is equivalent to $m$.
In a vacuum, the potential energy of a magnetic field per unit of volume is $\frac{d U}{d V}=\frac{B^{2}}{2 \mu_{0}}$ which is equivalent to $\frac{d E}{d V}=\frac{1}{2} \rho v^{2}$, therefore $\rho$ is equivalent to $\frac{1}{\mu_{0}}$, so the added mass is $\frac{\frac{8 R^{3}}{3 \mu_{0}}}{\frac{1}{\mu_{0}}}=\frac{8 R^{3}}{3} \rho$ Since only half of the object is immersed in water:
$m_{\text {added }}=\frac{4 R^{3}}{3} \rho$

Just like in method 1 , this can be seen as a mass of $M_{\text {added }}=\frac{4}{3} \rho R^{3}$ added to the cylinder in an inelastic collision, we can write the conservation of momentum equation
$\left(M_{\text {cylinder }}+M_{\text {added }}\right) \cdot u_{\text {after }}=M_{\text {cylinder }} \cdot u_{\text {before }}$
$M_{\text {cylinder }}=\pi \rho_{\text {cylinder }} \cdot(R)^{2} \cdot H=\frac{\frac{4}{3} \rho R^{3} u_{\text {after }}}{u_{\text {before }}-u_{\text {after }}}$
$\rho_{\text {cylinder }}=\frac{4 \rho R u_{\text {after }}}{3 \pi\left(u_{\text {before }}-u_{\text {after }}\right) \cdot(H)}$

## Taking measurements

I used a pixel ruler and vlc media player to measure the distance of the cylinder from the top of the screen, I took 5 measurements each frame and plotted the average in Desmos.
The disk hits the water in measurement 20 so we need to make two trend lines for before and after the impact.

| measure 1 [ px$]$ | measure 2 [ px$]$ | measure 3 [ px$]$ | measure 4 [ px$]$ | measure 5 [ $p x$ ] | avarage [px] | n [frame] |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 80 | 80 | 80 | 80 | 80 | 80 | 1 |
| 109 | 110 | 113 | 115 | 114 | 112.2 | 2 |
| 144 | 137 | 144 | 142 | 143 | 142 | 3 |
| 171 | 167 | 174 | 176 | 171 | 171.8 | 4 |
| 203 | 200 | 204 | 209 | 203 | 203.8 | 5 |
| 235 | 231 | 237 | 230 | 234 | 233.4 | 6 |
| 267 | 262 | 267 | 266 | 285 | 285.4 | 7 |
| 298 | 295 | 299 | 296 | 289 | 295.4 | 8 |
| 330 | 329 | 331 | 323 | 319 | 326.4 | 9 |
| 361 | 362 | 364 | 354 | 355 | 359.2 | 10 |
| 394 | 394 | 397 | 390 | 385 | 392 | 11 |
| 427 | 430 | 430 | 424 | 420 | 426.2 | 12 |
| 460 | 467 | 484 | 482 | 469 | 484.4 | 13 |
| 492 | 496 | 495 | 495 | 496 | 484.8 | 14 |
| 523 | 527 | 529 | 531 | 525 | 527 | 15 |
| 557 | 568 | 580 | 555 | 557 | 559 | 16 |
| 589 | 598 | 593 | 577 | 591 | 589.6 | 17 |
| 627 | 632 | 627 | 601 | 625 | 622.4 | 18 |
| 649 | 653 | 651 | 631 | 657 | 648.2 | 19 |
| 686 | 665 | 668 | 682 | 653 | 682.4 | 20 |
| 683 | 679 | 680 | 683 | 677 | 680.4 | 21 |
| 700 | 696 | 698 | 699 | 697 | 698 | 22 |
| 716 | 711 | 715 | 718 | 719 | 715.8 | 23 |
| 736 | 729 | 735 | 736 | 744 | 736 | 24 |
| 755 | 745 | 754 | 754 | 761 | 753.8 | 25 |
| 774 | 763 | 773 | 773 | 782 | 773 | 26 |
| 791 | 783 | 791 | 791 | 804 | 792 | 27 |
| 811 | 802 | 810 | 810 | 822 | 811 | 28 |
| 830 | 821 | 834 | 825 | 839 | 829.8 | 29 |
| 848 | 838 | 850 | 844 | 859 | 847.8 | 30 |
| 888 | 858 | 868 | 862 | 878 | 886.8 | 31 |
| 887 | 878 | 890 | 880 | 896 | 886.2 | 32 |
| 907 | 896 | 906 | 900 | 911 | 904 | 33 |
| 926 | 915 | 926 | 922 | 930 | 923.8 | 34 |
| 945 | 934 | 947 | 840 | 947 | 942.6 | 35 |
| 963 | 952 | 965 | 960 | 969 | 961.8 | 36 |
| 981 | 972 | 983 | 978 | 983 | 979.4 | 37 |
| 1000 | 989 | 1001 | 999 | 999 | 997.6 | 38 |
| 1017 | 1008 | 1018 | 1017 | 1018 | 1015.6 | 39 |
| 1035 | 1031 | 1034 | 1033 | 1036 | 1033.8 | 40 |
| 1051 | 1048 | 1049 | 1049 | 1050 | 1049.4 | 41 |
| 1069 | 1066 | 1068 | 1068 | 1067 | 1087.6 | 42 |
| 1079 | 1076 | 1078 | 1078 | 1075 | 1077.2 | 43 |
| 1087 | 1083 | 1091 | 1090 | 1080 | 1086.2 | 44 |
| 1094 | 1092 | 1099 | 1099 | 1092 | 1095.2 | 45 |
| 1103 | 1100 | 1107 | 1110 | 1100 | 1104 | 46 |
| 1112 | 1108 | 1118 | 1118 | 1113 | 1113.4 | 47 |
| 1123 | 1118 | 1125 | 1125 | 1120 | 1122.2 | 48 |
| 1130 | 1126 | 1132 | 1129 | 1130 | 1129.4 | 49 |
| 1137 | 1134 | 1142 | 1142 | 1135 | 1138 | 50 |
| 1143 | 1142 | 1148 | 1149 | 1144 | 1145.2 | 51 |
| 1150 | 1148 | 1156 | 1156 | 1152 | 1152.4 | 52 |
| 1180 | 1157 | 1184 | 1162 | 1159 | 1160.4 | 53 |



Since we only care about the first few frames after the disk hits the water (after a while instabilities develop and the system becomes hard to analyze), it is okay that the second line only matches the first few points


$u_{\text {after }}=18.5 \pm 0.4[p x /$ frame $]$
I measured:
$k=331[p x]$ (radius of the disk)
$H=265[p x]$ (height of the disk)
$\rho_{\text {water }}=1000\left[\mathrm{~kg} / \mathrm{m}^{3}\right]$
and finally:
$\rho_{\text {cylinder }}=\frac{4 \cdot(1000) \cdot 331 \cdot 18.5}{3 \cdot \pi \cdot(32.0-18.5) \cdot(265)}=73 \cdot 10^{1}\left[\mathrm{~kg} / \mathrm{m}^{3}\right]$
$\rho_{\text {cylinder }}=73 \cdot 10^{1} \pm 5 \cdot 10^{1}\left[\mathrm{~kg} / \mathrm{m}^{3}\right]$
This result is surprising since we don't see the disk float back up at the end of the video, but considering the fact we see very little change in the velocity when the disk is underwater, it is possible that eventually it rose back to the top (even the sudden change at $n=43$ is duo to the bottom of the disk hitting the edge of the screen forcing me to measure the height from the top, and most likely optic effects caused the top part to appear slower, or perhaps it was an actual rotation around the $x$ axis)

